

# 1 Rupture directivity of microearthquake sequences 2 near Parkfield, California

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5 [1] The direction of propagation is an important factor that  
6 affects the pattern of ground motion generated by an  
7 earthquake. Characterizing factors favoring a potential  
8 rupture propagation direction is thus an important task.  
9 Here we analyze the earthquake directivity of repeating  
10 earthquake sequences located on the San Andreas fault  
11 near Parkfield, California. All earthquakes of a sequence  
12 have very similar waveforms and have overlapping surface  
13 ruptures. We show that subtle variations of the transfer  
14 function between earthquakes of a common sequence can  
15 be interpreted as a change of apparent rupture duration.  
16 Relative apparent rupture durations are computed for all  
17 pairs of events at all available stations and for each  
18 sequence. We invert these measurements to obtain an  
19 estimation of the apparent rupture duration for each  
20 individual event of the sequence relative to a reference  
21 event. Variation of apparent rupture duration with azimuth  
22 attests for the rupture directivity. We show that the  
23 majority of analyzed microearthquakes presents a rupture  
24 in the south-east direction. We also show that, on a given  
25 repeating sequence, most earthquakes tend to show the  
26 same rupture direction. **Citation:** Lengliné, O., and J.-L. Got  
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## 30 1. Introduction

31 [2] Earthquake rupture is characterized, among other  
32 features, by its direction of propagation. This feature has  
33 important consequences in terms of potential damages as  
34 most of the energy will be carried out in the direction of  
35 rupture [e.g., *Boatwright*, 2007]. It is not yet clear which are  
36 the important parameters controlling the direction of rupture.  
37 As an example, the 1966 Parkfield earthquake has an  
38 inferred rupture propagation direction towards the south-  
39 east whereas the 2004 shock which ruptured the same fault  
40 patch propagated towards the north-west [*Bakun et al.*,  
41 2005]. Numerical models of dynamic rupture suggest that  
42 material contrast across the fault plane might induce a  
43 preferential rupture direction [e.g., *Andrews and Ben-Zion*,  
44 1997]. However there is still a lack of clear, direct, obser-  
45 vational evidence of a statistical preferential rupture direc-  
46 tion. Indeed, pre-stress on the fault plane is likely to be one  
47 of the factors controlling the rupture propagation direction.  
48 In order to uncover a preferential direction, one has to deal  
49 with a sufficient number of earthquakes to reduce the sta-

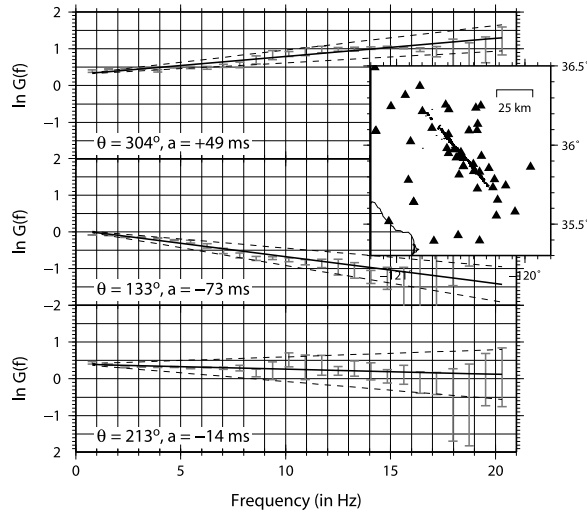
tistical noise induced by the effect of the pre-stress. Large 55  
earthquake datasets are mostly composed of low-magnitude 56  
events for which source characteristics are not accurately 57  
inferred. Here, we take advantage of repeating earthquake 58  
sequences previously isolated by *Lengliné and Marsan* 59  
[2009] to analyze the changes in directivity among earth- 60  
quakes showing similar waveforms, and to provide a statis- 61  
tical evidence of factors controlling the rupture directivity. 62  
Similar attempts have been recently conducted by *Kane et* 63  
*al.* [2009] and E. Wang and A. M. Rubin (Rupture direc- 64  
tivity of microearthquakes on the San Andreas fault from 65  
spectral ratio inversion, submitted to *Geophysical Journal* 66  
*International*, 2010). Repeating earthquake sequences used 67  
in this study have been identified based on (i) coherence 68  
criterion -coherence is a frequency dependent measure of 69  
similarity between waveforms-, (ii) nearly similar event 70  
magnitude and (iii) superposition of the source areas. The 71  
high number of events allows us to investigate the source 72  
process of multiple earthquake ruptures on the time span 73  
covered by the dataset (~22 years). Extracting significant 74  
information from these microearthquake sequences requires 75  
an adequate processing that makes use of the earthquake 76  
similarity. We employ a spectral ratio method which takes 77  
full advantage of the common ray paths of earthquakes of a 78  
common sequence to obtain precise estimates of their rela- 79  
tive sources parameters. Despite the extreme similarity of 80  
the waveforms, small variations are observed and can be 81  
exploited in order to indicate changes in the source process. 82  
Such source parameters are extracted from an inversion 83  
procedure that is devoted to incorporate precise information 84  
concerning the various forms of uncertainties arising in our 85  
problem. This processing provides us with relative apparent 86  
durations with confidence intervals for each earthquake of a 87  
sequence. A simple model of rupture allows us to interpret 88  
our results in terms of propagation direction. Our study aims 89  
at i) analyzing whether earthquakes occurring at the same 90  
location always have the same directivity or not, ii) detect- 91  
ing whether microearthquakes in the Parkfield area show a 92  
statistical preferential rupture direction or not. 93

## 2. Data Processing 94

[3] We use 334 repeating sequences, identified by 95  
*Lengliné and Marsan* [2009], totaling 2414 earthquakes 96  
with magnitude ranging from  $M_l = 1.0$  to 3.2. We follow the 97  
approach presented by *Got and Fréchet* [1993] to obtain the 98  
variation of rupture duration for a pair of earthquakes. We 99  
use 2.56 s-long P-wave records on the vertical component 100  
of short period stations of the Northern California Seismic 101  
Network (NCSN). All stations have a 100 Hz sampling 102  
frequency. We define  $N_{eq}$  as the number of earthquakes in 103  
the analyzed repeating sequence and  $n_{sta}$  the number of 104

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**Figure 1.** Variation of  $\ln G$  as a function of the frequency (in Hertz) for a pair of earthquake at three different stations. The errorbars in gray denotes the values of  $\ln G$  with their uncertainties ( $2\sigma$  confidence interval). The dark lines show the best linear fits and the dashed lines indicate the uncertainties, at the  $2\sigma$  level, on the slope determination. The value of the slope,  $a$ , as well as the azimuth of each station relative to the doublet barycentre are shown in the bottom left corner for each station. The enclosed figure is a map centered on the studied area representing earthquakes used in this study (black dots) and stations which recorded at least 100 pairs of earthquakes (black triangles).

105 stations which recorded at least two earthquakes of this  
 106 repeating sequence. We call  $x_i^k(t)$  the record of the  $i$ th  
 107 earthquake at station  $k$  ( $i \in [1; N_{eq}]$  and  $k \in [1; n_{sta}]$ ). For all  
 108  $n_{sta}$  stations we compute the modulus of the transfer func-  
 109 tion relating the Fourier transform of all possible pairs of  
 110 events. This is the modulus of the Wiener filter existing  
 111 between these two events. It is the least square estimate of  
 112 what is often called the “spectral ratio” between two events.  
 113 We call  $G_{ij}^k$  the modulus of this transfer function linking  
 114 signals  $x_i^k$  and  $x_j^k$ .  $G_{ij}^k$  is computed at frequency,  $f$ , by

$$G_{ij}^k(f) = \frac{|X_i^k(f)X_j^{*k}(f)|}{|X_j^k(f)X_i^{*k}(f)|}, \quad (1)$$

115 where  $X_i^k(f)$  is the Fourier transform of  $x_i^k$ , the star denotes  
 116 the complex conjugate,  $|z|$  is the modulus of  $z$  and the  
 117 overbar designates smoothed quantity. The two signals are  
 118 first iteratively aligned during the time-delay computation,  
 119 using cross-spectral analysis. We used a 1.28 s-long Tukey  
 120 tapering window; spectral densities are smoothed with the  
 121 Fourier transform of a Hann window of order two. The order  
 122 controls the smoothing width. The coherency, measuring the  
 123 similarity between the two signals at a given frequency is  
 124 given by

$$C_{ij}^k(f) = \frac{|X_i^k(f)X_j^{*k}(f)|}{\sqrt{|X_i^k(f)X_i^{*k}(f)|}\sqrt{|X_j^k(f)X_j^{*k}(f)|}}. \quad (2)$$

A mean coherency,  $\widehat{C}_{i,j}^k$ , is computed between 3 and 20 Hz. 125  
 The estimates of  $G_{ij}^k$  are kept when  $\widehat{C}_{i,j}^k$  is larger than 90%. 126

### 3. Model

[4] Let us consider *Brune's* [1970]  $f^2$  source which de- 128  
 scribes the frequency content for a kinematic fault model. In 129  
 such a model, the logarithm of the spectral ratio between 130  
 two earthquakes with corner frequency  $f_{c1}$  and  $f_{c2}$  can be 131  
 expressed as 132

$$\ln[G(f)] = \alpha + \ln \frac{1 + \left(\frac{f}{f_{c2}}\right)^2}{1 + \left(\frac{f}{f_{c1}}\right)^2}, \quad (3)$$

where  $\alpha$  denotes the logarithm of the seismic moment ratio 133  
 of the two events. We approximate the slope of  $\ln(G)$  134  
 computed at  $f = f_c$ , where  $f_c \simeq f_{c1} \simeq f_{c2}$  as earthquakes have 135  
 nearly similar sizes, with the slope of  $\ln(G)$  in the frequency 136  
 range  $[3-20]Hz$ . This approximation is relevant as  $\ln(G)$  is 137  
 quasi-linear in the frequency range below  $f_c$ . Following *Got* 138  
 and *Fréchet* [1993], the slope of  $\ln(G)$  can be approximated 139  
 at  $f = f_c$  as  $-\Delta f_c / f_c^2$ , where  $\Delta f_c = f_{c2} - f_{c1}$ . Assuming  $\tau \propto 1/f_c$ , 140  
 i.e., the rupture duration,  $\tau$ , is inversely proportional to the 141  
 corner frequency, we obtain that the slope of the logarithm 142  
 of the spectral ratio is proportional to the variation of rupture 143  
 duration. Therefore, taking the logarithm of  $G_{ij}^k(f)$  and 144  
 computing its slope with respect to frequency provides us 145  
 with an estimate of the apparent variation of rupture duration 146  
 between earthquakes  $i$  and  $j$  at station  $k$ . The slope of  $\ln$  147  
 $[G(f)]$  is computed with a simple least square fit where the 148  
 uncertainty  $\sigma_{ij}^k(f)$  on  $\ln(G_{ij}^k(f))$  is approximated by the stan- 149  
 dard deviation of a Gaussian distribution, and with 150

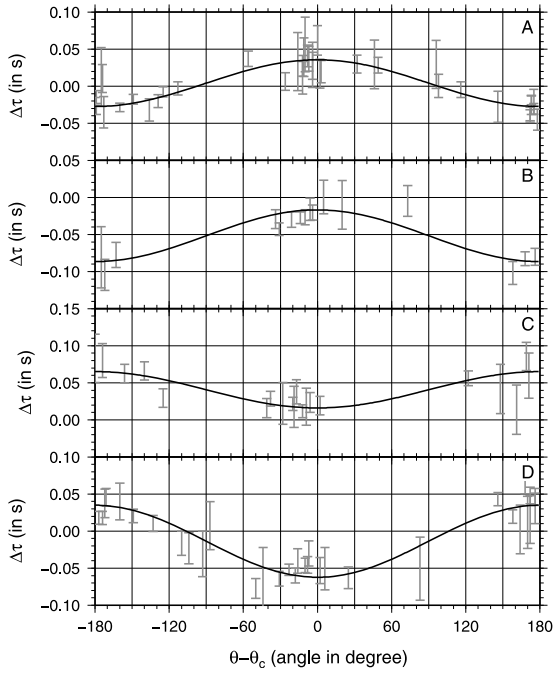
$$\sigma_{ij}^k(f) = \begin{cases} \frac{1 - (C_{ij}^k(f))^2}{(C_{ij}^k(f))^2} & \text{if } C_{ij}^k(f) > 0.9 \\ \infty & \text{else} \end{cases} \quad (4)$$

we also impose  $\sigma(f)$  to never be lower than 0.005 (equiv- 151  
 alent to  $C_{ij}^k(f) > 0.9975$ ) in order to not set unrealistically 152  
 small uncertainties in the case of very coherent waveforms. 153  
 The slope of  $\ln(G_{ij}^k(f))$  is denoted  $a_{ij}^k$  and its uncertainty 154  
 is  $\sigma_{a,ij}^k$ . We show in Figure 1 the typical variation of  $\ln(G(f))$  155  
 for a pair of earthquakes at three different stations. We observe 156  
 a clear linear decay whose fit provides values of  $a$ . It demon- 157  
 strates that although waveforms are very similar, variations 158  
 exist among them and can be analyzed. 159

[5] For all possible pairs of earthquakes at all possible 160  
 stations, we use this method to obtain estimates of the 161  
 variation of rupture duration. As we have measurements for 162  
 all possible pairs, and all measurements have to be coherent 163  
 between them, we can write a system of linear equations in 164  
 order to estimate the apparent rupture duration  $\Delta\tau_i^k$  for all 165  
 events relative to the first event of the sequence. The 166  
 problem we need to solve is linear and can be written as 167

$$\mathbf{d} = \mathbf{G}\mathbf{m}. \quad (5)$$

The data vector,  $\mathbf{d}$ , is composed by  $a_{ij}^k$  values. The parameter 168  
 vector,  $\mathbf{m}$ , is made up of the  $\tau_i^k$  values which are the 169



**Figure 2.** Variation of apparent rupture duration,  $\Delta\tau_i^k$  as a function of the azimuth  $\theta - \theta_c$  for several earthquakes, for different sequences. The error bars in gray denote the  $2\sigma$  confidence interval of  $\Delta\tau_i^k$  and the dark curve is the best cosine fit. We distinguish the 4 cases A–D. In cases A and D, values of  $\Delta\tau$  are both positive and negative and thus correspond to the model presented in (10). For cases B and C, values of  $\Delta\tau_i^k$  are either entirely positive or entirely negative and which corresponds to the model represented by equation (9).

170 apparent rupture duration of event  $i$  at station  $k$  that we want  
171 to determine. As  $a_{ij}^k = \tau_j^k - \tau_i^k$ , the Jacobian matrix  $\mathbf{G}$  only  
172 comprises 0, +1 and -1. Solution to equation (5) is provided  
by

$$\tilde{\mathbf{m}} = (\mathbf{G}'\mathbf{C}_D^{-1}\mathbf{G} + \mathbf{C}_M^{-1})^{-1} (\mathbf{G}'\mathbf{C}_D^{-1}\mathbf{d}_{obs} + \mathbf{C}_M^{-1}\mathbf{m}_{prior}), \quad (6)$$

173 where  $\tilde{\mathbf{m}}$  is the *a posteriori* parameter vector and  $\mathbf{m}_{prior}$  is  
174 the *a priori* parameter vector,  $\mathbf{G}'$  is the transpose of  $\mathbf{G}$   
175 [Tarantola, 2005]. The data covariance matrix is  $\mathbf{C}_D$  and the  
176 *a priori* model covariance matrix is  $\mathbf{C}_M$ .  $\mathbf{C}_D$  is non empty  
177 only on the main diagonal with all  $\sigma_{a,ij}^k$  values. We assign an  
178 *a priori* parameter uncertainty of 1 s except for the first  
179 event for which we assign a very small *a priori* uncertainty.  
180 All *a priori* parameters are set to 0 s. By fixing a very  
181 small *a priori* uncertainty on  $m_{prior}(1)$ , we thus impose that  
182  $\tilde{m}(1) \sim 0$  and thus that all results will be relative to  $m(1)$ . The  
183 *a posteriori* uncertainties are obtained with

$$\tilde{\mathbf{C}}_M = (\mathbf{G}'\mathbf{C}_D^{-1}\mathbf{G} + \mathbf{C}_M^{-1})^{-1}. \quad (7)$$

184 We finally obtain the apparent durations of rupture  $\Delta\tau_i^k$  for  
185 each event of the processed sequence and for each available  
186 station. All these estimates are relative to the apparent  
187 duration of the first event of the sequence, chosen as the

reference event. In order to keep only well resolved relative  
188 rupture duration estimates we discard all *a posteriori* para-  
189 meters with associated uncertainties greater than 0.05 s.  
190

## 4. Results

191

[6] For a kinematic source model, with a rupture propa-  
192 gating horizontally at velocity  $v_r$  along the fault strike, the  
193 apparent rupture duration  $\tau_r$  is given by Haskell [1964]  
194

$$\tau_r = \frac{L}{v_r} \left( 1 - \frac{v_r}{c} \sin \phi \cos \theta \right), \quad (8)$$

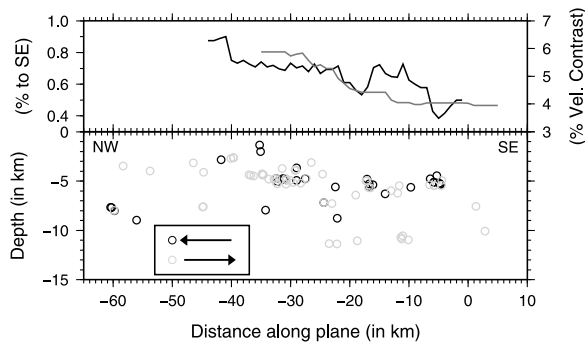
where  $c$  is the  $P$ -wave velocity,  $\theta$  is the azimuth of the  
195 station relative to the rupture direction and  $\phi$  is the take-off  
196 angle. The distance  $L$  corresponds to the distance over  
197 which the rupture propagates;  $L$  equals the total fault plane  
198 length in the case of a unilateral rupture. As we are dealing  
199 with relative measurements, our results comprise both  
200 source properties not only of the earthquake  $i$ , but also of the  
201 first earthquake of the sequence used as a reference. We  
202 make the hypothesis that the rupture velocity for two  
203 earthquakes of a same sequence is similar. We also suppose  
204 that the rupture process of both earthquakes takes place on a  
205 fault plane with the same orientation and the same rupture  
206 mechanism. This is suggested from the focal mechanisms of  
207 earthquakes in the area which are almost entirely strike-slip  
208 [Thurber *et al.*, 2006]. We define  $\theta_c = 140^\circ$  as the azimuth  
209 of the San-Andreas fault at Parkfield in the south-east  
210 direction [Thurber *et al.*, 2006]. For each earthquake of a  
211 sequence we want to determine its direction of rupture. The  
212 rupture direction is defined here as the direction for which  
213 the rupture propagates over the longest distance. This distance  
214 is equal to the fault plane length in a purely unilateral  
215 rupture and might be as small as  $L/2$  for a perfectly bilateral  
216 rupture. Two scenarios are considered: i) both earthquake  $i$   
217 and the reference earthquake have the same rupture direction  
218 or ii) the two earthquakes have opposite rupture directions.  
219 From equation (8) and considering that the apparent rupture  
220 duration is the difference between the initiating phase and  
221 the last stopping phase we can compute the relative apparent  
222 rupture duration. Depending on the two proposed cases, the  
223 relative apparent rupture duration will be respectively  
224

$$\Delta\tau_i^k = \frac{L_i - L_0}{v_r} \left( 1 \pm \frac{v_r}{c} \sin \phi_i^k \cos \theta'^k \right), \quad (9)$$

$$\Delta\tau_i^k = \frac{L_i - L_0}{v_r} \pm \frac{L_i + L_0}{c} \sin \phi_i^k \cos \theta'^k, \quad (10)$$

where  $\theta' = \theta - \theta_c$ . The lengths  $L_i$  and  $L_0$  represent the  
225 distance over which the rupture propagates for earthquake  $i$   
226 and the reference earthquake respectively. Distinguishing  
227 whether the ruptures we are inferring are closer to the uni-  
228 lateral case than to the bilateral case would require com-  
229 paring  $L_i - L_0$  with the fault plane length. As this last  
230 measurement is not known precisely we do not differentiate  
231 between these two cases and only investigate the direction  
232 of rupture as defined previously. We fit the azimuthal varia-  
233 tion of  $\Delta\tau_i^k$  with a function of the form  
234

$$\Delta\tau_i^k = A_i + B_i \cos(\theta'^k) \sin(\phi^k), \quad (11)$$



**Figure 3.** (bottom) Location of the 95 sequences, along the San Andreas fault plane, each of them showing at least one direction of rupture (gray circle: predominant rupture towards the south-east, black circle: towards the north-west). The horizontal axis is the distance along fault and its origin is defined as the hypocenter of the 2004,  $M_w = 6$  mainshock. The vertical axis is the depth. (top) Proportion of sequences with a dominant rupture direction to the south-east (black). The proportion is computed from along strike bins of 10 km length when at least seven sequences fall into the considered bin. Average velocity contrast along the San Andreas fault plane from values by *Zhao et al.* [2010] (gray line).

235 where  $A_i$  and  $B_i$  are the parameters to be determined,  $\sigma_A$  and  
 236  $\sigma_B$  are their corresponding standard deviations. Such a  
 237 function represents a valid fit for both scenarios (equations (9)  
 238 and (10)). These two scenarios can be distinguished based on  
 239 the signs of  $\Delta\tau_i^k$  values. As  $\frac{v_c}{c} < 1$ , (equation (9)) shows that  
 240  $\Delta\tau_i^k$  values may be positive or negative but can not be both. In  
 241 this case,  $|A| > |B|$  and both earthquakes (reference and tested  
 242 earthquake) rupture propagates in the same direction. The  
 243 rupture direction is obtained from the sign dependence of  
 244 equation (9). Rupture propagates in the direction of  $a_c$  if the  
 245 sign is positive, in the opposite direction else. It results that  
 246 the direction of rupture of the tested earthquake is given by  
 247 the azimuth for which  $|\Delta\tau_i^k|$  is minimum. When  $|B| > |A|$ ,  
 248  $\Delta\tau_i^k$  takes positive and negative values. The rupture direction  
 249 of the event  $i$  is controlled by the sign dependence of equation  
 250 (10) and thus by the sign of  $B$ . This rupture direction is given  
 251 by the azimuth for which  $\Delta\tau_i^k$  is minimum. We note that our  
 252 two scenarios prescribed the rupture to occur in two directions  
 253 only: on the direction of  $\theta_c$  or opposite to it.

254 [7] The relative apparent rupture duration presents (Figure 2)  
 255 a clear azimuthal pattern that is well fitted by the proposed  
 256 cosine form (equation (11)). Our model (equations (9) and  
 257 (10)) implies that the variation of  $\Delta\tau$  with azimuth is solely  
 258 explained by the difference in location of hypocenters,  
 259 eventually leading to changes in rupture direction. We may  
 260 wonder if any other possible change between the two earth-  
 261 quakes can also modify the proposed patterns. As proposed  
 262 by *Got and Fréchet* [1993], we can first exclude a change of  
 263 attenuation as it induces only a weak variation of  $\Delta\tau$  com-  
 264 pared to the one observed. These authors also showed that a  
 265 change of rupture velocity or a change of focal mechanism  
 266 due to local variations of the fault plane geometry will not  
 267 produce a pattern similar to the one proposed (equations (9) or  
 268 (10)) and thus will be discarded in the following analysis due  
 269 to the resulting high misfit with equation (11).

[8] In order to avoid interpreting fits which are not well  
 constrained and to reject ambiguous cases, we reject esti-  
 mates of rupture direction when  $\sigma_B > 5 \cdot 10^{-3} s$  and when  
 $|B| < 2\sigma_B$ . We finally obtain 95 sequences for which at least  
 one rupture direction has been determined. These 95 se-  
 quences provided 273 estimates of rupture directions, 188 of  
 which are in the direction of  $\theta_c$ , i.e., to the southeast which  
 represents 69% of all estimates. Restricting our analysis  
 with sequences comprising at least 3 estimates of rupture  
 direction, we find 35 sequences with a total of 197 rupture  
 direction estimates, 135 of them (or 69%) being oriented  
 toward the southeast. We can thus infer that micro-  
 earthquakes in our dataset preferentially rupture in the  
 southeast direction. We divided the 35 sequences, with at  
 least 3 directivity estimates, based on the most abundant  
 rupture direction of each sequence. We obtain 26 sequences  
 with a dominant directivity towards the southeast and 9  
 sequences with a dominant directivity towards the north-  
 west. We also investigate whether the direction of rupture  
 varies for earthquakes in a common sequence or not. For the  
 35 sequences with at least 3 direction estimates, 84% of the  
 ruptures on a sequence are found in the same direction. This  
 suggests that earthquakes on an identified repeating source  
 tend to have the same direction of rupture. We also show in  
 Figure 3 the repartition along the fault plane of all the 95  
 sequences with their preferential rupture direction. We  
 observe a decrease of the preferential direction of rupture  
 towards the south-eastern bound of the fault segment.

## 5. Discussion and Conclusion

[9] Several mechanisms might be invoked in order to  
 explain the preferential rupture direction of the earthquake  
 towards the south-east. One of them involves material  
 contrast across the fault plane. The rupture on such bima-  
 terial interface is influenced by normal stress reduction in a  
 favored direction which produces the directivity effect. Such  
 a bimaterial model may represent an appropriate description  
 of our studied zone. Indeed we analyze earthquake se-  
 quences located on the San Andreas fault which is supposed  
 to mark an important material contrast [e.g., *Thurber et al.*,  
 2006; *Zhao et al.*, 2010]. Numerical models and theoretical  
 studies suggest that the earthquake rupture directivity is  
 preferentially oriented in the slip-direction of the less-rigid  
 material (towards the south-east) [e.g., *Weertman*, 1980;  
*Andrews and Ben-Zion*, 1997; *Ben-Zion and Andrews*,  
 1998; *Cochard and Rice*, 2000; *Rubin and Ampuero*,  
 2007; *Ampuero and Ben-Zion*, 2008]. It has however been  
 proposed by *Harris and Day* [2005] that the propagation  
 direction might not be a direct consequence of the material  
 contrast as pre-stress can also influence the rupture direction  
 [e.g., *Andrews and Harris*, 2005; *Ampuero and Ben-Zion*,  
 2008] and bilateral rupture are also found in numerical  
 rupture on bimaterial interface [*Harris and Day*, 1997]. A  
 suggestion for a preferred rupture direction however comes  
 from the observation of an asymmetric distribution of  
 immediate aftershocks of microearthquakes on the San  
 Andreas fault plane [*Rubin and Gillard*, 2000]. Our results  
 indicate that microearthquakes in the Parkfield area statis-  
 tically show a preferential rupture direction, i.e., a system-  
 atic tendency of the moment density distribution relative to  
 the hypocenter to skew toward the southeast. Such an  
 asymmetry has also been evidenced by *Wang and Rubin*

331 (submitted manuscript, 2010). Furthermore, this preferred  
 332 direction of rupture is in agreement with the direction pre-  
 333 dicted from the velocity contrast across the fault plane. Our  
 334 results are also supported by the progressive variation of the  
 335 velocity contrast along the fault plane. As evidenced by  
 336 *Thurber et al.* [2006] and *Zhao et al.* [2010] the material  
 337 contrast is weaker towards the southeast portion of the fault  
 338 segment near the 2004, mainshock location. We see on  
 339 Figure 3 that this reduction of the velocity contrast, imaged  
 340 by *Zhao et al.* [2010], closely follows the decrease of the  
 341 proportion of earthquake sequences showing a preferential  
 342 direction towards the south-east. Due to the averaging  
 343 procedure used to estimate the velocity contrast, only the  
 344 variation of the velocity contrast should be considered not  
 345 the absolute values. Our findings suggest that material  
 346 contrast across the fault plane is a possible cause inducing  
 347 this statistical preferential rupture direction. This effect is  
 348 revealed only after the analysis of a sufficient number of  
 349 similar earthquakes. Other factors, as the variability of the  
 350 pre-stress along the fault plane -which may randomly affect  
 351 the rupture direction of an individual earthquake- are  
 352 reduced by the statistical averaging. This is also evidenced  
 353 at the scale of the asperity for sequences with a sufficient  
 354 number of events. At this scale, we observe that earthquakes  
 355 on a common asperity show a statistically preferential rup-  
 356 ture direction. It suggests that, at the asperity scale as well,  
 357 the rupture direction is influenced by material contrast and is  
 358 also dependent on other effects as pre-stress at the source  
 359 location. We note however that if the amplitude of stress  
 360 heterogeneities is scale-dependent, the microearthquake  
 361 observations presented in this study might be hard to  
 362 extrapolate to large earthquakes.

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